



Precessional and half-precessional climate forcing of Mid-Devonian monsoon-like dynamics

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Abstract. A Devonian magnetic susceptibility (MS) record obtained on limestones ranging from the Uppermost-Eifelian to the Lower-Givetian and located on the southern border of the Dinant Synclinorium in Belgium was selected for time-series analysis. In these carbonate ramp and platform deposits, spectral analyses highlight persistent, high-frequency cycles in both the MS-signal and the microfacies curve, reflecting environmental and climate changes. These metre-scale variations in the MS-signal are interpreted as changes in the flux of magnetic minerals towards the marine system, most likely controlled by monsoon rainfall-intensity. By combining chrono- and biostratigraphic information with theoretical knowledge of sedimentation rates in different depositional environments, these cycles are interpreted as astronomically driven and dominated by precession. It is hypothesized that during precession minima (longitude of the perihelion $\approx 270^\circ$) the trans-equatorial pressure gradient reaches a maximum and intensifies the northeasterly monsoonal circulation. The consequent increase in moisture transport towards the continent leads to enhanced precipitation and runoff, which, in turn, induces a higher flux of detrital material – including magnetic minerals responsible for the MS-signal – towards the marine system. Moreover, this unique high-resolution climate signal reveals half-precessional cycles. These cycles suggest that during precession maxima (longitude of the perihelion $\approx 90^\circ$) southeasterly monsoonal circulation strengthened due to an increased inter-hemispheric pressure gradient.

1 Introduction

It is only since the late seventies that variations in the Earth's orbit were accepted as the pacemaker of the Quaternary ice ages (Hays et al., 1976). To leave the sphere of controversy, the theory needed several well-established feedback mechanisms and assumptions about the nonlinear responses of ice sheets to the applied astronomical forcings. In an extreme hot, greenhouse world, without any major ice sheets, ice-related feedback mechanisms and nonlinearities do no longer apply. Consequently, other mechanisms must explain the amplification of subtle variations in the Earth's orbital parameters into global climate and environmental changes. Except for the important positive water-vapour feedback mechanisms (Rind et al., 1991; Flohn et al., 1992) or the vegetation and soil feedback mechanism (Kutzbach et al., 1996; Dekker et al., 2007), a clear understanding of how this amplification occurred in an ice-free hot greenhouse world is lacking.

The Greenhouse Devonian Period (418–361 Ma; Kaufmann, 2006) is characterised by climate boundary conditions that differ completely from those of the Quaternary: atmospheric CO₂ concentration was up to 10 times higher compared to today's value (Bernier, 2006), continents were concentrated in the Southern Hemisphere (SH), global temperatures and sealevel were high and vast shallow epicontinental seas allowed reefs to develop at high latitudes, reaching 65° N–55° S during the Middle-Devonian (Copper, 2002). Moreover, the Devonian is a key period in the evolution of life on Earth with, for example, the fish-tetrapod transition occurring during the early Middle-Devonian (Niedzwiedzki

et al., 2010) and the invasion of the continents by land plants. The latter development influenced climate significantly by altering continental albedo, roughness, thermal properties, potential evaporation and CO₂ consumption (Le Hir et al., 2011). Therewithal, enhanced pedogenesis resulted in intensified continental weathering and export of its breakdown-products to seas and oceans (Berner, 1997; Algeo et al., 2001). In this study, the mathematical techniques of time-series analysis were applied to a well-documented 85 m-thick Eifelian/Givetian (Middle-Devonian) section from Southern Belgium (Mabille and Boulvain, 2007; Boulvain et al., 2010). The studied section, characterised by a high accumulation rate and high sampling resolution, contains a climate signal that is highly suitable to study the influence of Milanković (eccentricity, obliquity and precession) and sub-Milanković cycles on the Middle-Devonian climate of the Rhenohercynian Basin. The main objective of this study is to reveal the finger print of these cycles in the section, in order to obtain a better insight into the response to astronomical forcing of an extremely hot Greenhouse climate.

2 Geological setting

The Eifelian-Givetian “La Couvinoise” quarry is located along the southern flank of the Dinant Synclinorium, in southern Belgium (Fig. 1), which is part of the Rhenohercynian fold-and-thrust belt. The studied section corresponds to a mixed carbonate-siliciclastic ramp, located in the Rhenohercynian basin, on the southern margin of the Avalonian microcontinent (Cocks et al., 1997). This basin was limited by the Euramerican Old Red Continent to the north and by the Normannian and Mid-German highs to the south (Ziegler, 1990). The apparent polar wander paths of Avalonia and Euramerica are well established (Torsvik et al., 1993; Mac Nio-caill and Smethurst, 1994). Torsvik et al. (1993) estimated the paleolatitude of Southern Britain at 389 Ma to be 6.9° S. During the Middle-Devonian, the studied section and Southern Britain were only a few hundred kilometres apart (Cocks et al., 1997) and consequently this value serves as a good estimate for the paleolatitude of the section.

The section has been subdivided in different microfacies corresponding to two main depositional settings: open marine at the base of the section and fore-reef at the top (Fig. 2; Mabille and Boulvain, 2007). Open marine facies correspond to a succession of argillaceous mudstones and packstones carbonates with crinoids and brachiopods. Rudstones with stromatoporoids debris, tabulate and rugose corals, followed by packstones and grainstones with peloids characterise the overlying fore-reef carbonate facies (Mabille and Boulvain, 2007).

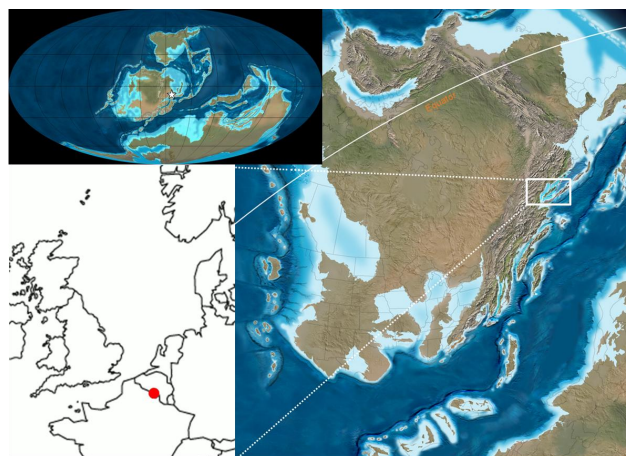


Fig. 1. Mid-Devonian paleogeography (after Ron Blakey, Northern Arizona University Geology) and present-day situation of the “La Couvinoise” section.

3 Materials and methods

3.1 Magnetic Susceptibility as a climate proxy

Magnetic Susceptibility (MS) represents the degree of magnetization of a material in response to an applied magnetic field. It quantitatively measures the amount of magnetic minerals in a sample. The MS of a sediment forms a proxy for the rate of supply of the magnetic susceptible lithogenic or detrital fraction to the marine system (Ellwood et al., 2000). The MS measurements were made on a KLY-3 Kappabridge device. Three measurements were made on each sample and the average MS was taken to construct the record. Samples were weighted with a precession of 0.01 g (detailed description of methodology in Da Silva and Boulvain, 2006). In the studied section, large-scale MS variations can be ascribed to relative sea-level changes (Mabille and Boulvain, 2007; Da Silva et al., 2009). On a shorter time-scale, climate, through the intensity of wind and precipitation, plays a decisive role in respectively determining the eolian and riverine flux of magnetic minerals to the marine system (Ellwood et al., 2000; Hladil et al., 2006; Riquier et al., 2010) as it was, in this case, the Rheic Ocean (Fig. 1). For the Eifelian and Givetian in the southern border of the Dinant Synclinorium, diagenesis did not obliterate the primary signal, as the MS remains strongly related to depositional environment. Also, an estimate of detrital quartz content in the thin section demonstrates a strong correlation between the abundance of detrital quartz and MS values (Fig. 9 in Mabille and Boulvain, 2007). While the detrital quartz does not carry the MS signal, it is a good indicator of the detrital input. It appears, thus, that MS is likely correlated to detrital input and not entirely controlled by diagenesis (Mabille and Boulvain, 2007). Moreover, MS revealed itself as a powerful tool to establish accurate stratigraphic correlations between the studied section

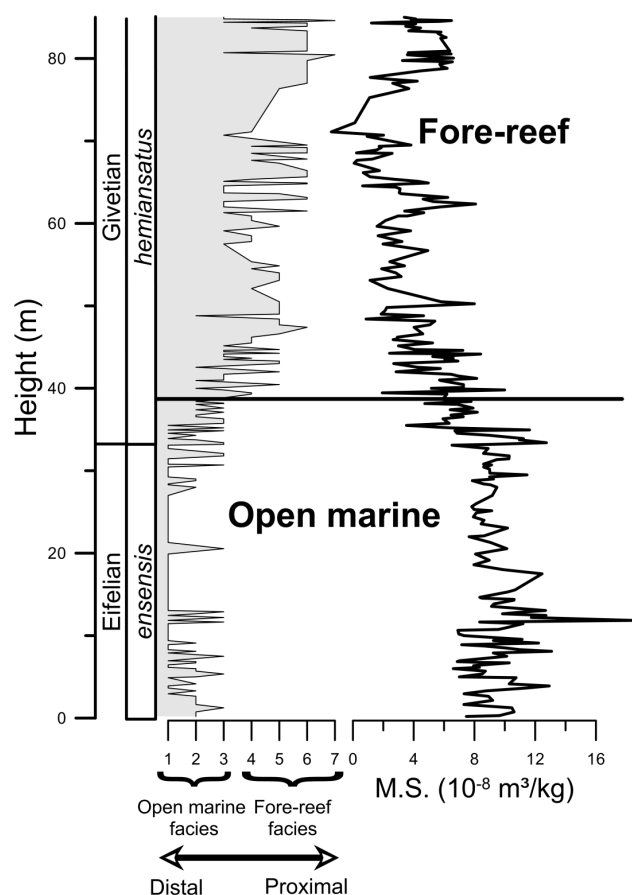


Fig. 2. The “La Couvinoise” section Magnetic Susceptibility and facies belts (modified after Boulvain et al., 2010). Key to facies numbers: open marine facies: 1 = Argillaceous mudstone; 2 = Argillaceous bioclastic wackestone; 3 = Brachiopods-crinoids argillaceous packstones; Fore-reef facies: 4 = Floatstones and rudstones with reefal debris; 5 = Stromatoporoids bindstones; 6 = Peloidal packstones-grainstones; 7 = Peloidal-crinoidal grainstones.

and a contemporaneous section, a few kilometres to the west, which is characterised by a major divergence in terms of sedimentary dynamics (Mabille and Boulvain, 2007). Therefore, the MS signal used for time-series analyses can be considered primary and robust to lateral sedimentological and facies variations.

In today’s world, the response of different tropical monsoon systems to orbital, Milanković forcing is clearly demonstrated (e.g., Kutzbach, 1981; Kutzbach and Liu, 1997; Tuenter et al., 2003). In comparison, the Devonian tropical climate potentially exhibited an even more intense monsoonal circulation (Streel et al., 2000), characterised by seasonally wet-and-dry climates (Cecil, 1990). As early as 1916, Joseph Barrell interpreted the red sandstones, shales and conglomerates of the Devonian system in the British Isles as “*fluvatile in origin; laid down over river flood plains*

by streams in times of flood, exposed to air in times of drought. They record in this way the existence of an alternation of seasons of rainfall and drought - a climate with an arid season, but not an arid climate. [...] Sediment was brought into these basins by rivers from the bordering uplands and from the more distant regions to the northwest. The excess beyond what was laid down by the rivers in time of flood to maintain their grade across the sinking basins was carried through to the shallow sea which lay on the surface of the continent to the southwest”. As the studied section is located in this shallow sea of the Rhenohercynian basin, the flux of detrital material into this depositional environment will reproduce the seasonal pattern. Ever since Barrell’s interpretation (1916), paleoclimatic indicators confirmed this view (e.g., Woodrow et al., 1973; Heckel and Witzke, 1979; Woodrow, 1985; Witzke and Heckel, 1988; Sageman et al., 2003; Cressler, 2006; Cressler et al., 2010), and a consensus on the seasonal wet-and-dry climate model for the Devonian of the Euramerican coastal plains exists. The paleolatitude combined with the proximity of large continental areas (as well Euramerica as Gondwana) play an important role in determining the potential for an intense monsoonal circulation over Euramerica (Witzke, 1990). Given the significant influence of astronomical forcing on modern monsoonal systems, it is highly probable that also during the Devonian period, Milanković cycles influenced the climatic factors controlling the detrital inputs to the Rhenohercynian basin (i.e., wind and precipitation) and/or of carbonate productivity, both affecting the MS signal (Riquier et al., 2010).

3.2 Time-series analysis procedure

The frequency composition of the magnetic susceptibility signal is analysed via spectral analysis, using the multitaper method (MTM; Thomson, 1982), as implemented in the SSA-MTM Toolkit (Ghil et al., 2002). Before the implementation of MTM spectral analysis, long-term trends were removed from the MS data by subtracting the continuous linear trend. A constant spacing (in depth) between sample points in the series was generated by linear interpolation at intervals approximating the average sample spacing. The MTM-method was performed using 3 tapers to compromise between resolution and sidelobe reduction. Small variations in the accumulation rate behave like phase modulations and introduce erroneous spectral peaks (Muller and MacDonald, 2000). The MTM-method is selected because it averages these sidelobes into the main peak and, thereby, gives a superior estimate of the true spectral power. Noise in paleoclimate signals is frequently “1/f noise” or “red noise”, which means that the power spectrum of the background shows a strong enhancement at the low frequency end (Muller and MacDonald, 2000). The 90 % and 99 % confidence limits (CLs) are calculated to visualize the most significant deviations from the null hypothesis of red noise (robust AR(1) estimation, median smoothing window width = $(10\Delta t)^{-1}$, Log Fit, $f_{Nyquist} = 2/\Delta t$).

The microfacies curve is obtained by digitizing the microfacies log, resulting in square wave data. The step-like changes in microfacies can only be represented using multiple additional sine and cosine components during a Fourier transformation, polluting the spectrum (Proakis and Menolakis, 1996; Weedon, 2003). Therefore, spectral analysis on the microfacies curve was carried out using the Walsh and Blackman-Tukey (BT) method for spectral analysis. Walsh power spectra are based on the use of square waves instead of sine and cosine waves. The equivalent of a Fourier Transform can be generated using the Walsh transform (Beauchamp, 1975; Weedon, 1989). The Blackman-Tukey method is based on the Fourier transform as part of the autocovariance sequence. The autocovariance sequence contains the same frequency oscillations as the microfacies series, but is no longer characterised by step-like changes.

Frequency-selective filters or band-pass filters isolate a specific range of frequencies from a signal. Consequently, it is an excellent device to assess the behaviour of a specific range of frequencies in a studied signal. The AnalyseSeries software is used to calculate the filtered signals (Paillard et al., 1996).

The analysis of amplitude modulations (AM) constitutes a useful tool in cyclostratigraphy: the astronomical theory states that the amplitude of the precession parameter is modulated by eccentricity. To examine possible amplitude modulations in the studied signal, the precession parameter was first extracted from the signal by band-pass filtering. Subsequently, its amplitude envelope was extracted by complex demodulation, based on the Hilbert transform (Bloomfield, 2004).

3.3 Possible distortions of the astronomical frequencies

When interpreting cyclic sedimentary alternations as a consequence of periodic changes in the orbital parameters of the Earth, one should be aware of the fact that the assumption of a 1:1 recording of astronomical signals in carbonate sediments is questionable (Westphal et al., 2004). The original astronomical signal can be altered by the climatic system, responding in a nonlinear way to the applied forcing. Moreover, a sediment column recording climatic changes never represents these temporal changes in a linear manner. For example, variations in sedimentation rate, bioturbation, differential compaction and pressure dissolution can be responsible for distortions of the orbital frequencies in sedimentary records. To account for these possible overprinting factors, some assumptions and restrictions must be respected. On carbonate platforms and ramps, sediment production rate is above other factors, a function of carbonate productivity. The main trend is a decrease in the carbonate productivity with increasing depth, with maximum carbonate production in the reefal environments (Da Silva et al., 2009). Therefore, the two main sedimentary environments represented in the “La Couvinoise” section (Fig. 2) are

characterised by significantly different accumulation rates. A significantly higher accumulation rate characterises the fore-reef environment compared to the open marine environment (Da Silva et al., 2009). In depositional environments with different accumulation rates, cycles of equal duration will be expressed by sediment packages of different thickness. This phenomenon could render the application of spectral analysis on a distance-series including different depositional settings impossible. However, if one assumes that accumulation rate within the same sedimentary environment is semi-constant, MTM-spectral analysis can be carried out separately for the two main sedimentary environments of the “La Couvinoise” section. For the “La Couvinoise” section, no indications for hiatuses or for differential compaction within a sedimentary environment are found.

4 Results and discussion

4.1 Recognition and identification of astronomical cycles

For the open marine environment, the Walsh and BT spectral plots of the microfacies data are characterised by an elevated spectral power, around frequency 0.7 (Fig. 3a). Moreover, an increased spectral power that extends above the 90 %CL can be found at the same frequency on the MTM spectral plot of the MS data. In the fore-reef environment, the Walsh and BT spectral plots of the microfacies-data (Fig. 3b) exhibit three significant spectral peaks: around frequency 0.05 m^{-1} , around 0.3 m^{-1} and around 0.7 m^{-1} . Remarkably, the MTM spectral plot of the MS data reveals the exact same pattern (Fig. 3b). A similar spectrum for magnetic susceptibility and microfacies indicates that microfacies and MS record the same environmental variations. Therefore, common indications for astronomical forcing are more credible, whereas differences between both proxies call for a critical evaluation of the extent to which astronomical forcing influenced environmental changes.

The only time-constraint available in the studied section is the existing biostratigraphy and more precisely the delimitation of the *ensensis* conodont biozone between the bottom of the section and 33.75 m. The *hemiansatus* conodont zone cannot be used as a time-constraint because its top does not occur in the “La Couvinoise” section. In well-studied sections from the Eifel region, the *ensensis* biozone covers about 8 % of the duration of the Eifelian (Weddige, 1977; Bultynck et al., 1988; Kaufmann, 2006). As the Eifelian lasted for $3.8 \pm 6 \text{ Myr}$ (Kaufmann, 2006), the *ensensis* biozone has a duration of $304 \pm 480 \text{ kyr}$. Therefore, one of the 23 observed cycles in the *ensensis* biozone (1.25–1.66 m thick) lasted for $13.2 \pm 20.9 \text{ kyr}$. Despite the huge spread in these values, the observed persistent cyclicity can most probably be assigned to the $\sim 18\text{-kyr}$ Middle-Devonian precession cycle (with quasiperiods of 16 824 and 19 886 kyr; Berger and

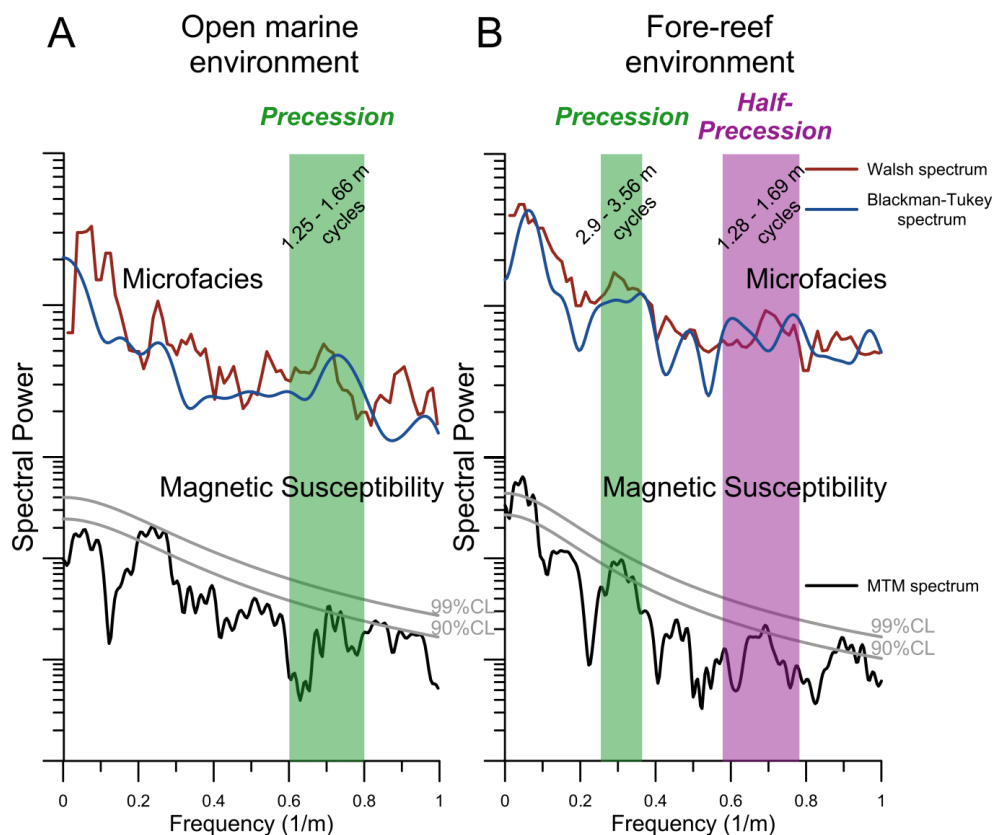


Fig. 3. Spectral analyses on the different depositional environments of the “La Couvinoise” section separately. **(A):** spectral analyses on the microfacies series and Magnetic Susceptibility series of the lower part of the section (open marine environment). **(B):** spectral analyses on the microfacies series and Magnetic Susceptibility series of the upper part of the section (fore-reef environment). CL = Confidence Limit.

Loutre, 1989; Berger et al., 1992). One can note that the ~ 33 kyr period of the obliquity cycle lies at the uppermost end of the error bar, but it can almost certainly be excluded as a driving force of the observed cycles since the influence of obliquity on seasonal insolation at low latitudes is very small compared to the precession effect and ice-related teleconnections to higher latitudes are negligible. The interpretation of the distinguished cyclicity as precession cycles implies a duration of ~ 414 kyr for the *ensensis* biozone and an astronomically calibrated average accumulation rate of 8.15 cm kyr^{-1} for this biozone in the studied section. Taking into account the superior sedimentation rate in a fore-reef depositional setting compared to an open-marine environment, it is expected that sedimentary cycles related to precession are thicker in the fore-reef environment, compared to the open marine environment (Da Silva et al., 2009). Therefore, in the upper part of the section, the 2.9–3.56 m cycles are interpreted as the result of precession. The 1.28–1.69 m cycles in this part of the section have a period that is about half the precessional period. If this interpretation is correct, the average sedimentation rate in the fore-reef part of the section lies around 18 cm kyr^{-1} . The astronomically derived sedimentation rates (8.15 cm kyr^{-1} in the open marine en-

vironment and 18 cm kyr^{-1} in the fore-reef environment) are in the same order of magnitude as the average sedimentation rate calculated for the whole Givetian for the paleolocation of Southern Belgium (i.e., 10.5 cm kyr^{-1} ; Boulvain et al., 2010) or as the estimates of accumulation rate made for a contemporaneous (Mid-Eifelian to basal-Givetian) section in the nearby Eifel region (Rhenish Massif): 12.53 cm ka^{-1} (Weddige, 1977; Bultynck et al., 1988; Kaufmann, 2006). The two latter average sedimentation rates, extracted from literature, were calculated over different depositional environments, ranging from open marine to internal platform. The fact that these average values fall right in the middle between our astronomically derived sedimentation rates for the open marine and fore-reef environment can be seen as a strong indication that the correct astronomical interpretation was made.

In the open marine environment 1.25–1.66 m cycles were interpreted as the imprint of precession, while in the fore-reef environment precession is depicted in the form of 2.9–3.56 m cycles. These interpretations render possible the transformation of the magnetic susceptibility distance-series into a time-series (Fig. 4). The precessional cycles are extracted from the distance-series by band-pass filtering the lower part of the

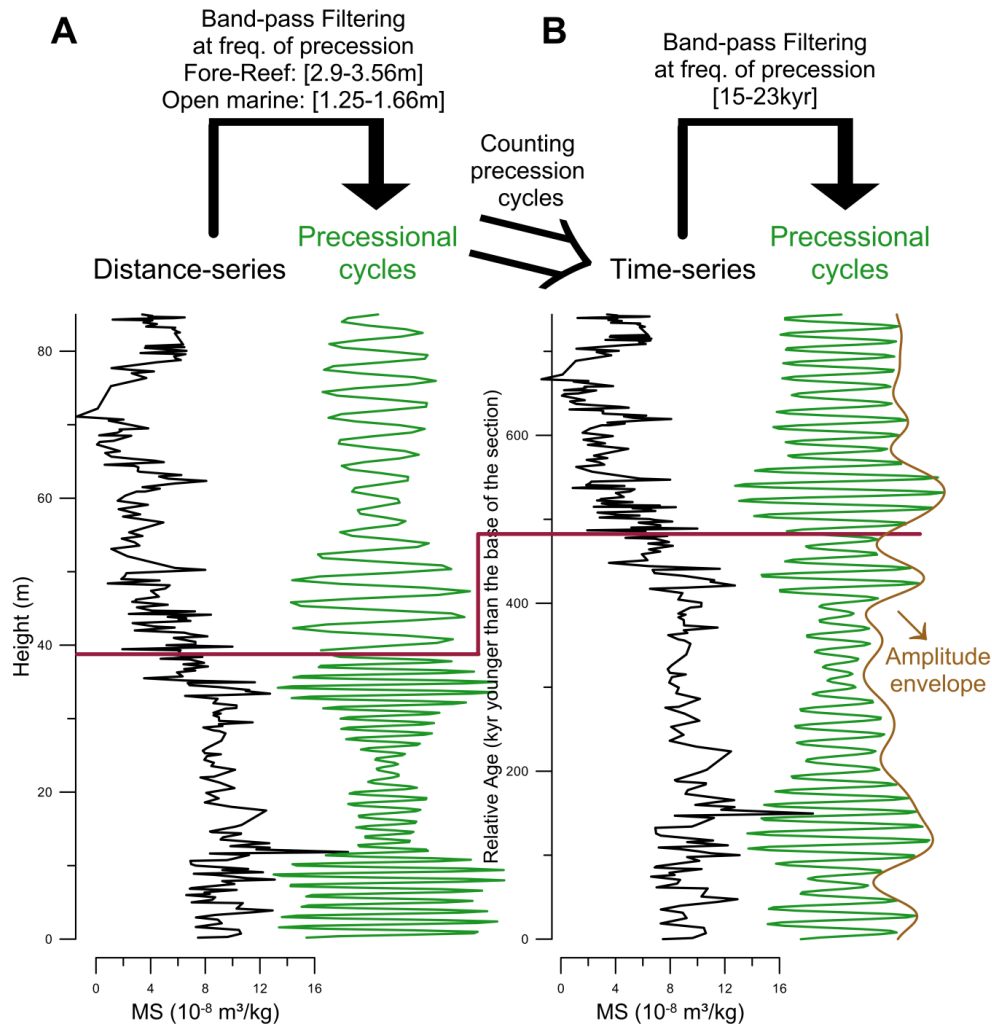


Fig. 4. From distance-series to time-series. **(A):** Magnetic Susceptibility series of the “La Couvinoise” section plotted against a distance-axis (black) and band-pass filtered Magnetic Susceptibility signal at the frequencies of precession (green, band-pass filtered between 1.25–1.66 m in the lower part and between 2.9–3.56 m in the upper part of the section). **(B):** Magnetic Susceptibility series of the “La Couvinoise” section plotted against a time-axis (black) and band-pass filtered Magnetic Susceptibility signal at the frequency of precession (band-pass filtered between 15 and 23 kyr; green) + amplitude envelope of the precessional signal (brown).

section between frequencies 0.6 and 0.8 m^{-1} (1.25–1.66 m periodicity) and by filtering the upper part of the section between frequencies 0.28 and 0.35 m^{-1} (2.9–3.56 m periodicity; Fig. 4a). By counting the precessional cycles and multiplying them by the period of precession (i.e., ~ 18 kyr), the MS-series can be rescaled relative to a time-axis (Fig. 4b). On Fig. 4, 41.3 precessional cycles can be counted in the “La Couvinoise” section. This implies that the section was deposited in ~ 744 kyr. When the constructed time-series is band-pass filtered at the frequency of precession (between 0.043 and 0.067 kyr^{-1} , i.e., 15–23 kyr periodicity), a ~ 100 -kyr cyclicity is observed in the amplitude envelope of the precessional cycles (Figs. 4b and 5a). This observation strongly suggests the amplitude modulation of the precessional cycles by eccentricity.

Using the constructed time-series (Fig. 4b) instead of the distance-series, spectral analysis is again carried out on the MS data (Fig. 5b). The reenactment of spectral analysis brings two significant improvements: first, the “La Couvinoise” section is considered as a whole, rather than in two different parts, so that the continuity of the cycles can be better tested. Second, the problem of accumulation rate variations within a depositional environment is bypassed. The resulting spectral plot (Fig. 5b) shows a very strong peak at the frequency of precession (17.5 kyr). Of course, the magnitude of this spectral peak is partly conditioned by the procedure of time-axis composition. For this reason, no further conclusions can be drawn from the magnitude of this spectral peak. However, there is no reason to doubt the statistical significance of this peak as the time-axis composition is based on

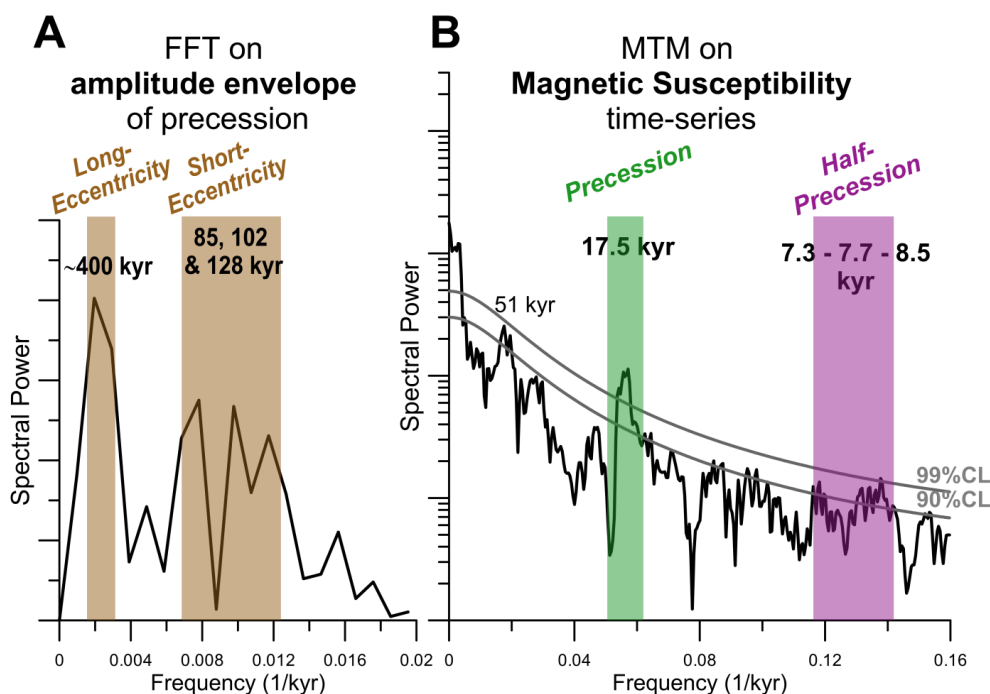


Fig. 5. Spectral analyses on the MS series of the “La Couvinoise” section as a whole. **(A):** fast Fourier Transform (FFT) periodogram of the amplitude envelope of the precessional signal. A linear “spectral power” axis is chosen to give an immediate sense of how much of the variance comes from each cycle. **(B):** multitaper method (MTM) power spectrum of the “La Couvinoise” MS data as a time-series. A logarithmic “spectral power” axis is chosen to make the statistical significance of peaks evident, particularly for small high-frequency peaks. CL = Confidence Limit.

a robust spectral analysis of the MS signals of the different sedimentary environments (Fig. 3). Also, Fig. 5b allows to explore the influence of other astronomical parameters than precession: one would expect a spectral peak around 0.03 and 0.01 kyr⁻¹ if respectively obliquity and 100-kyr eccentricity would influence the MS signal of the “La Couvinoise” section (Berger et al., 1992). At these frequencies no outstanding spectral peak can be found, advocating that obliquity and eccentricity do not strongly influence the MS signal in a direct way. To test for the indirect effect of eccentricity-forcing, Fig. 5a shows the spectral plot (Fast Fourier Transformation, FFT) of the amplitude envelope of the precession cycles that contains strong peaks at periodicities 128, 102 and 85 kyr. These peaks correspond well to the theoretical periods of short-eccentricity (i.e., 125 and 95 kyr). However, it should be noted that the double peak associated with short-eccentricity can also be the result of slightly varying sedimentation rates, behaving as phase modulations. Also in Fig. 5a, a low-frequency spectral peak occurs between frequency 0.002 and 0.003, which can be related to long 405-kyr eccentricity. These results demonstrate that the amplitude envelope of the precession signal is modulated by eccentricity. The recognition of the amplitude modulation of the precession cycles by eccentricity can be considered as an independent confirmation of the accuracy of the time-series construction.

Moreover, on Fig. 5b, another cluster of spectral peaks can be observed between frequencies 0.12 and 0.14 kyr⁻¹ (7.3–8.5 kyr periodicity). At periodicity 7.3 kyr the spectral peak even exceeds the 99 % CL. This pattern seems to suggest the presence of half-precessional cycles recorded in the MS stratigraphy of the “La Couvinoise” section.

In Fig. 6a, the microfacies curve is rescaled and plotted against a time-axis, based on the same procedure as for the MS series. Both the Walsh transform (Fig. 6b) as the Fourier transform (Fig. 6c) were used as the basis for calculating a band-pass filter centred on the frequency of precession (between 15 and 23 kyr). Both filters demonstrate that a large fraction of the microfacies variation in the studied section happens on a precessional time-scale. Moreover, an important amplitude modulation of the precessional signal becomes clear (Fig. 6b and c), characterised by a strong 93 to 128 kyr period (Fig. 7a). This amplitude modulation is interpreted as the influence of the short-eccentricity cycle. Analogous to the MS data, the constructed microfacies time-series (Fig. 6) is used instead of the distance-series to reenact spectral analysis on the “La Couvinoise” section as a whole. The resulting spectral plots (Fig. 7b) all exhibit a significant spectral peak at frequency 0.0547 kyr⁻¹ (18.3 kyr period) that can be related to precession. The Walsh spectrum shows additional precession-related peaks at periods of 14.3 and 16.1 kyr. In the frequency range, where half-precessional

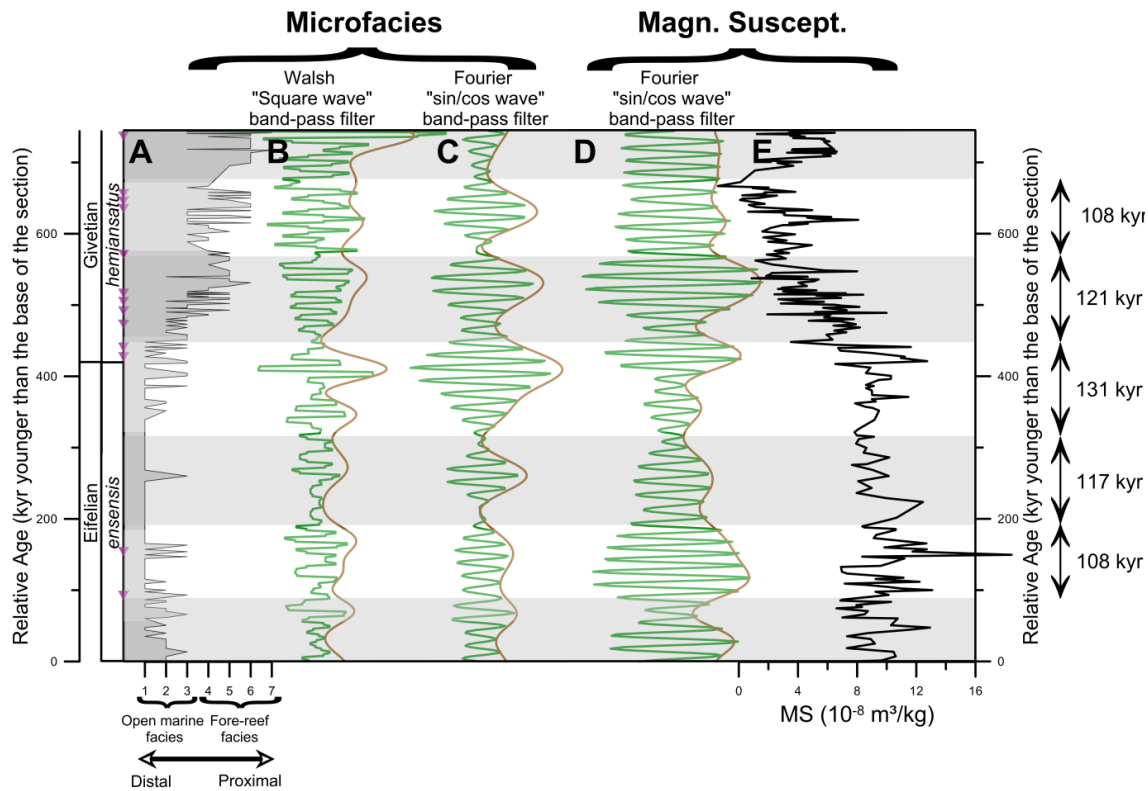


Fig. 6. (A): microfacies series of the “La Couvinoise” section plotted against a time-axis. (B): band-pass filtered microfacies signal at the frequency of precession (filtered between 15 and 23 kyr; green) based on the Walsh Transform + amplitude envelope (brown). (C): band-pass filtered microfacies signal at the frequency of precession (15–23 kyr; green) based on the Fourier Transform + amplitude envelope (brown). (D): band-pass filtered MS signal at the frequency of precession (15–23 kyr; green) based on the Fourier Transform + amplitude envelope (brown). (E): MS signal of the “La Couvinoise” section plotted against a time-axis. Purple triangles indicate those precessional cycles within which the half-precessional cycles are well developed.

cycles are expected, i.e., between 0.10 and 0.14 kyr^{-1} (between 7 and 10 kyr periodicity), the BT and Walsh spectrum suggest 3 or 4 spectral peaks. Although not very extensive, these peaks seem to confirm the presence of an important half-precessional cycle. On Fig. 6a, the stratigraphic positions where half-precessional cycles are well developed in the microfacies curve are indicated by purple triangles.

Obviously, the amplitude modulation for both MS and microfacies must exhibit a similar pattern throughout the section to justify the interpretation of these modulations as eccentricity-forced. Therefore, the MS precessional cycles and their amplitude envelope (Fig. 6d) are plotted next to the microfacies precessional cycles and their amplitude envelope (Fig. 6b and c). The comparison of the two amplitude envelopes shows that they indeed show a similar development throughout the section and gives rise to the delineation of five full short-eccentricity cycles. The period of these eccentricity cycles between 108 and 131 kyr is in quite good agreement to the theoretical periods of short-eccentricity (i.e., 95 and 125 kyr).

So far, the amplitude modulation of the precession signal by eccentricity is seen as a confirmation of the astronomical interpretation. However, it is equally crucial to test the robustness of the amplitude modulation itself. In order to test if the amplitude modulation depends on few extreme MS values, or if it is truly the expression of amplitude-changes of precession-related MS variations, following procedure was carried out. The magnetic susceptibility time-series consists of 236 points (Fig. 8a). The time-series was divided into 11 intervals of 20 points and an interval of 16 points. Of each interval, the maximal and minimal MS values were removed. In other words, 12 times 2 MS measurements were taken out of the time-series. The result is a time series with 212 points (i.e., 10 % less than the original time-series; Fig. 8b). This process, in which 12 times 2 points are deleted, was repeated to obtain a time-series with 188 points (i.e., 20 % less than the original time-series; Fig. 8c). In an analogous way, time-series consisting of 164 points (i.e., 30 % less than the original time-series; Fig. 8d) and 140 points (i.e., 40 % less than the original time-series; Fig. 8e) were created. The elimination of 10 % extreme points from the

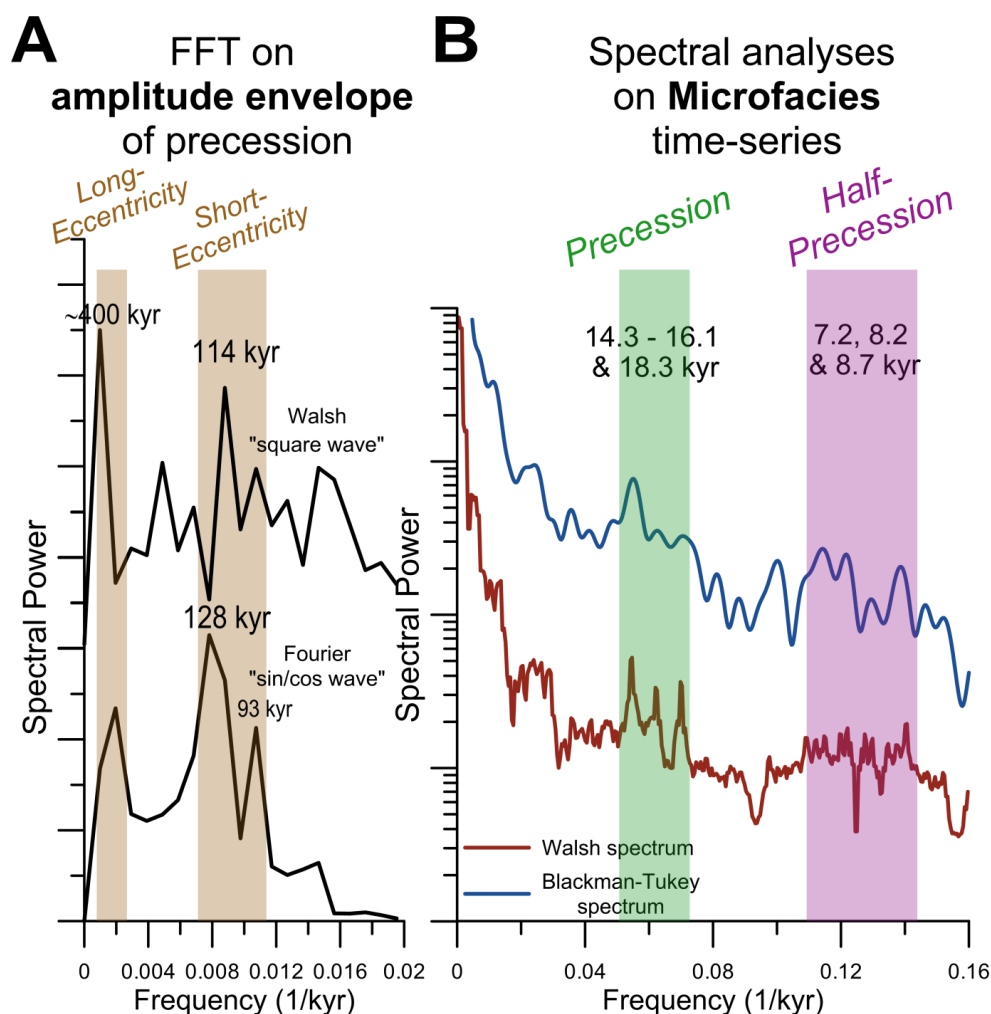


Fig. 7. Spectral analyses on the microfacies series of the “La Couvinoise” section as a whole. **(A):** FFT periodogram of the amplitude envelope of the precessional signal. Upper: based on Walsh Transform (Fig. 6b); Lower: based on Fourier Transform (Fig. 6c). A linear “spectral power” axis is chosen to give an immediate sense of how much of the variance comes from each cycle. **(B):** MTM power spectrum of the “La Couvinoise” microfacies curve as a time-series. A logarithmic “spectral power” axis is chosen to make the statistical significance of peaks evident, particularly for small high-frequency peaks. CL = Confidence Limit.

time-series influences the amplitude modulation to a limited extent, and would lead to a similar delineation of the 100-kyr eccentricity cycles. When 20 % or more extreme points are removed, amplitude modulation is hardly observable in parts of the section that are characterised by a relatively low sampling resolution. Still, in high-resolution intervals, the removal of extreme points does not obliterate the amplitude modulation pattern. Therefore, Fig. 8 clearly demonstrates that the amplitude modulation is not the result of a few extreme MS values, but of a true amplitude modulation of the precession-related cyclicity by eccentricity.

When the precessional variations in microfacies and MS are compared, it is clear that in the lower part of the section, these variations are in antiphase with each other (Fig. 9). In this part of the section, more proximal microfacies cor-

respond to MS minima. To the contrary, in the upper part of the section, the precessional variations in the microfacies curve are in phase with precessional variations in the MS curve and thus proximal microfacies correspond to MS maxima (Fig. 9). This observation can be explained by the evolution of mean MS values on relative proximity transects (Fig. 9c; Da Silva et al., 2009). In the lower part of the section, deposition occurred in an open marine environment. In such a setting, a shift to a more proximal environment is characterised by a higher carbonate productivity (diluting the MS signal) and water agitation (preventing the deposition of magnetic minerals). In the upper part of the section, deposition occurred in a fore-reef environment. Here, shifts to a more proximal environment usually encompass the occurrence of peloidal pack- and grainstones. These microfacies,

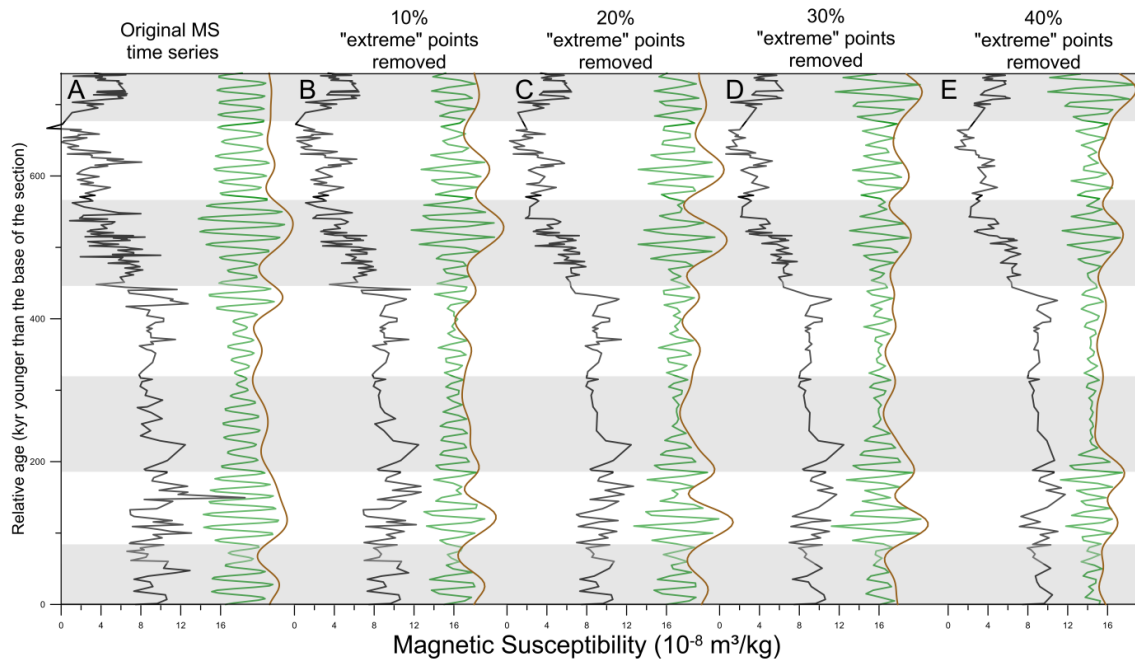


Fig. 8. Robustness-test of the amplitude modulation of the precessional signal in the MS time-series. **(A):** original MS time-series (black), its precessional signal (green) and the amplitude envelope of the precessional signal (brown). **(B, C, D, E):** MS time-series of which the 10 %–20 %–30 %–40 % most extreme MS values are removed (black) and its precessional signal (green).

which are indicative of a more restricted environment, are characterised by a higher MS compared to the more distal fore-reef microfacies (Fig. 9c; Da Silva et al., 2009).

4.2 Palaeoclimatological interpretation

The cycles driven by the astronomical parameters cannot be used as a geological metronome without an understanding of the way astronomical cycles influence the ferromagnetic minerals delivery processes towards the basin. The precessional forced MS variations in the “La Couvinoise” section are interpreted as the expression of variations in monsoon-circulation intensity. Here, it is hypothesized that the monsoon-circulation, rainfall intensity and the continental detrital sediment flux all depend on precession.

In Fig. 10, the monsoonal circulation on the east Euramerican continent is schematically illustrated and is based on the understanding of present monsoonal circulation. During Southern Hemispheric (SH) summer, a strong low-pressure cell develops above the Euramerican continent, allowing the Intertropical Convergence Zone (ITCZ) to cross the continent. Under such a configuration, the eastern part of the continent is dominated by a northeastern wind flow. During SH winter, the ITCZ zone is positioned in the Northern Hemisphere and the eastern part of the continent is dominated by the southeasterly trade winds.

Strong paleoclimate evidence from Africa, Australia, the Mediterranean and the Indian Ocean shows that today’s monsoonal circulation is intensely controlled by precession

(Clemens et al., 1991, 2010; deMenocal et al., 1993; Beaufort et al., 2010; Dickson et al., 2010; Ziegler et al., 2010). Most likely, similar mechanisms are responsible for the precessional cycles found in the studied “La Couvinoise” section. During a precession minimum (longitude of the perihelion $\approx 270^\circ$), the Earth is situated near the perihelion (close to the sun) during austral summer solstice. This astronomical configuration causes a significant increase in the amount of insolation that reaches the southern subtropics during SH summer. The additional heat results in an intensified rising of warm air from the surface of the Euramerican continent. Consequently, a significantly deeper low-pressure cell develops above the continent pushing the ITCZ southwards compared to a “normal” Southern Hemispheric Summer. Consequently, the transequatorial pressure gradient between the continental low-pressure cell and the subtropical high-pressure cell in the north increases and causes an intensified monsoonal circulation. In other words, the intensified zonal pressure gradient during a precession minimum enhances wind velocity and enables an increased moisture transport towards the continent, leading to enhanced precipitation. More precipitation leads to increased runoff that in turn leads to an increased flux of detrital material from the continent to the marine system and consequently, to a maximum in magnetic susceptibility. Therefore, precessional forcing of the amount of precipitation appears as an important forcing factor in the subtropical climate during the Devonian.

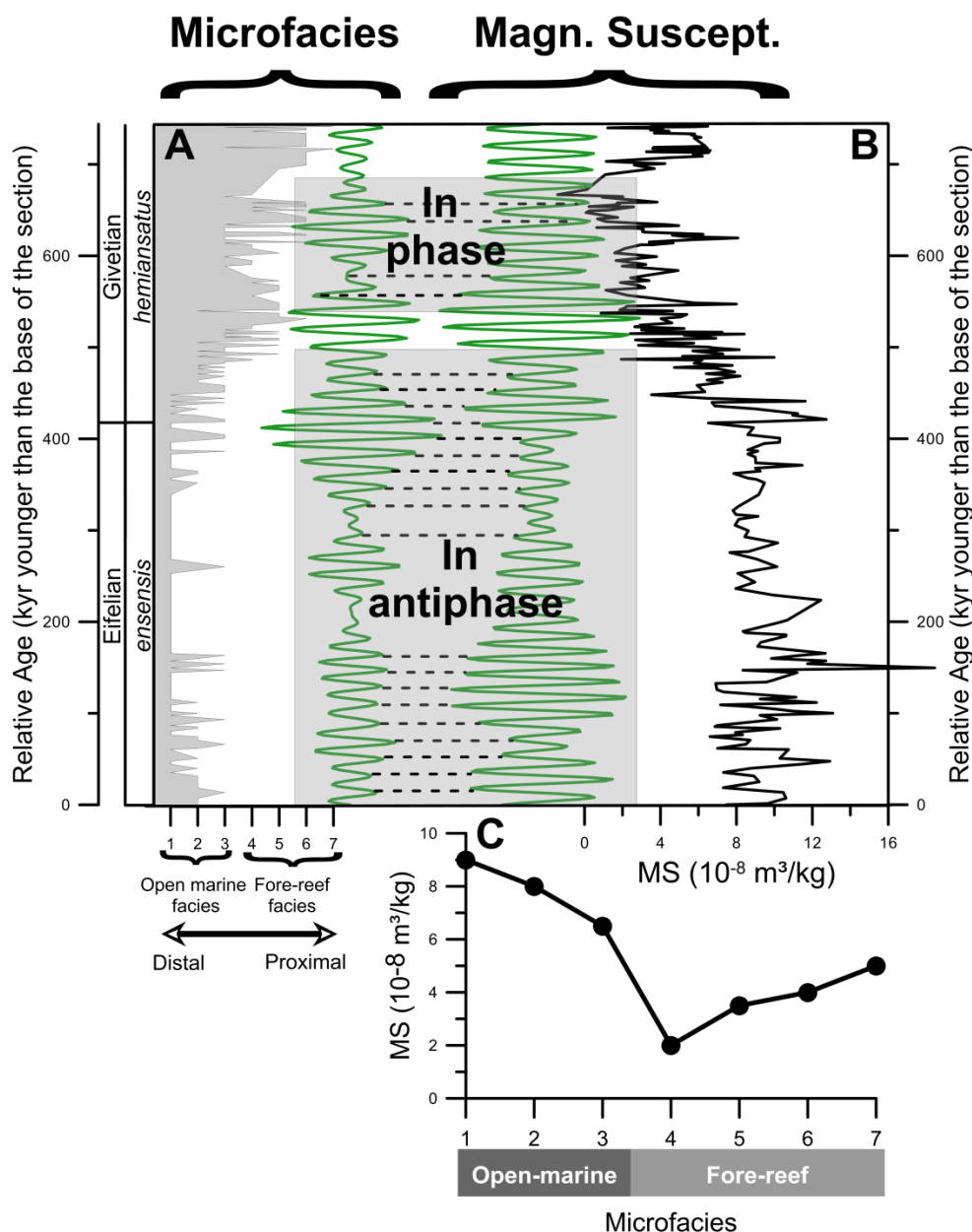


Fig. 9. Precessional microfacies variation vs. precessional Magnetic Susceptibility variation. (A): microfacies curve plotted against a time-axis and band-pass filtered microfacies curve at the frequency of precession (between 15 and 23 kyr; left). (B): Magnetic Susceptibility series plotted against a time-axis and band-pass filtered MS series at the frequency of precession (between 15 and 23 kyr; right). (C): evolution of mean MS values on a relative proximity transect (after Da Silva et al., 2009). The reversal of the slope of the MS evolution on a relative proximity transect explains why precession-scale variations of microfacies and MS are in antiphase in the lower part of the section and in phase in the upper part of the section.

However, the MS stratigraphy of the “La Couvinoise” section also revealed cyclicity at the frequency of half a precession cycle (~ 9.8 kyr). Berger et al. (2006) demonstrated the presence of an 11-kyr period in modern equatorial insolation. This sub-Milanković period, which is a harmonic of precession, is claimed to occur in paleoclimate archives of modern monsoonal systems (Hagelberg et al., 1994; Trauth

et al., 2003; Turney et al., 2004; Sun and Huang, 2006; Verschuren et al., 2009). Several mechanisms were suggested to explain how the 11-kyr astronomical cycle is transferred to climate: an amplified response of tropical precipitation and temperature to changes in maximum summer insolation in both hemispheres (Hagelberg et al., 1994), atmospheric and oceanic teleconnections (Turney et al., 2004), low-latitude

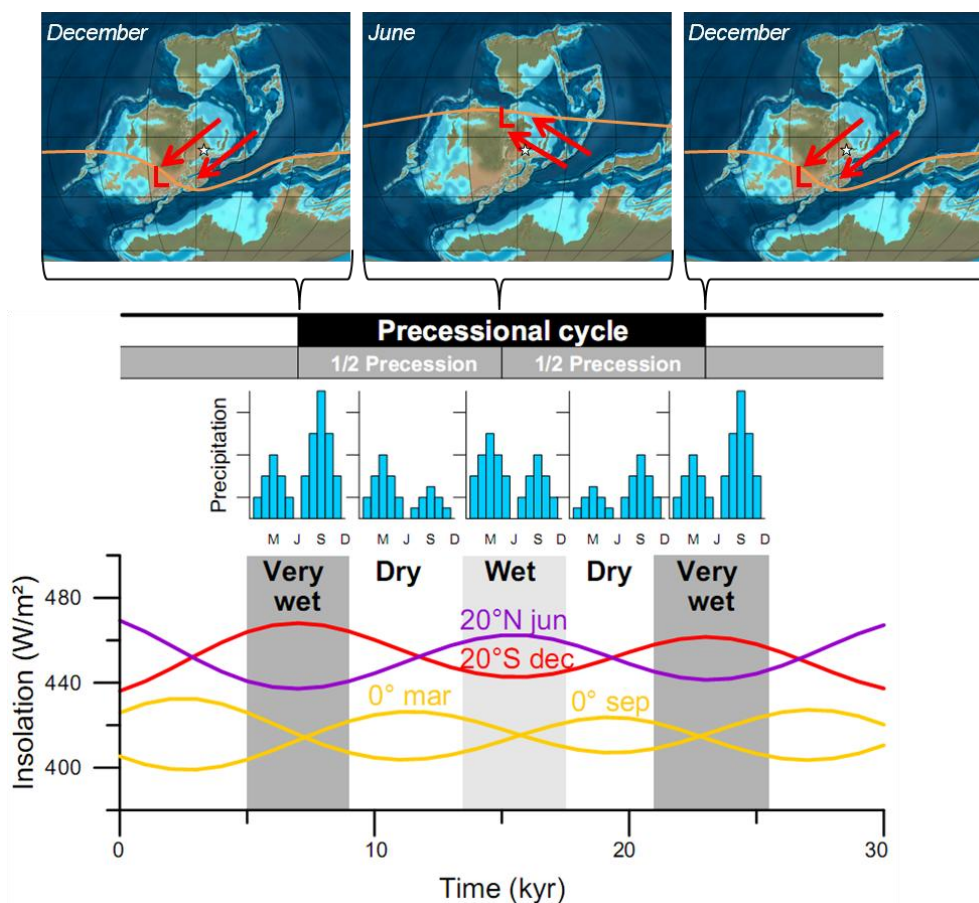


Fig. 10. Monsoon-like circulation in eastern Euramerica during the Mid-Devonian (Paleogeographic map from Ron Blakey, Northern Arizona University Geology). Intertropical Convergence Zone (ITCZ) in orange. Important trade winds are indicated by red arrows. L: Low pressure cell. A white star indicates paleogeographic position of the study site. The schematic representations of the monthly precipitation pattern illustrate the dependence of precipitation intensity on precession. Insolation pattern is based on Berger et al. (1992) and insolation strength is estimated considering a 5 % lower solar luminosity.

insolation forcing of monsoon-intensity (Short et al., 1991; Sun and Huang, 2006; Verschuren et al., 2009). Moreover, a coupled atmosphere/ocean/vegetation model demonstrated variations at periods of about 10 kyr in the monsoonal runoff, caused by the dynamic response of vegetation to precessional forcing (Tuenter et al., 2007). In this study, the observed half-precession cyclicity demonstrates far-reaching analogies with the half-precessional cycles found in lacustrine sediments of the past 25 000 yr in East-Africa (Verschuren et al., 2009). Both sedimentary archives are located in the ITCZ migration belt at almost the same latitude, both archives are positioned on the eastern coast of a large continent and, thus, influenced by a northeastern and southeastern monsoon. Consequently, the forcing mechanism of the half-precessional cycles in the studied Middle-Devonian section can be interpreted in a similar way. Figure 10 shows that every ~ 9 kyr the summer insolation peaks in either the northern or the southern tropics. As a consequence, every ~ 9 kyr, the inter-hemispheric insolation gradient is at a max-

imum and either the southeasterly or northeasterly monsoon is intensified. Moreover, low March or September insolation weakens the corresponding rainy season (Berger et al., 2006; Verschuren et al., 2009). In the tropical paleoenvironment of the studied section, this means that the flux of detrital material not only exhibits maxima during precession minima, but that it is also intensified during precession maxima. More detailed climate study, involving climate modelling, will be needed to fully understand the relationship between the half-precession period in equatorial insolation and its consequences in tropical and monsoonal climates.

5 Conclusions

Time series analyses of a high-resolution magnetic susceptibility signal from the Uppermost Eifelian and Lower Givetian of Southern Belgium indicates that astronomical forcing affects detrital input into the Rhenohercynian Basin. During

the Uppermost Eifelian and Lower Givetian, the studied “La Couvinoise” magnetic susceptibility signal exhibits the direct influence of a 9.8-kyr half-precessional cycle and an 18-kyr-precessional cycle. Obliquity and eccentricity do not seem to strongly influence the MS signal, at least in a direct way. However, the amplitude-envelope of the precessional signal in MS and microfossils clearly demonstrates a ~100-kyr variation that can be attributed to short 100-kyr eccentricity. Since this is the expected pattern according to astronomical theory, the interpretation of the observed cyclicity as precession-dominated can be considered reliable.

Over the last 3 Myr, low-latitude climate variability was closely intertwined with the onset and development of high-latitude continental ice-sheets, dominated by a ~41 kyr and ~100 kyr rhythm. In the studied Middle-Devonian climate record, these periodicities are completely absent. Therefore, it is suggested that in the Devonian extreme greenhouse world, without large continental ice sheets, millennial-scale tropical climate variability is dominated by precession and is not (or to a very limited extent) linked to high-latitude climate variability. Most probably, changes in precession affected the transequatorial pressure gradient, its consequent monsoonal dynamics, rainfall intensity and the input of detrital material in the Rheic Ocean. The recognition of a half-precessional cycle suggests the importance of the inter-hemispheric pressure gradient in the steering of monsoonal intensity. The role of the big Euramerican continent that stretches along both sides of the equator is undoubtedly very important in the development of a double monsoon-intensity maximum per precession cycle.

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